

## 05. Turing Machines and Spacetime.

1. Types of Spacetimes
2. Classical Spacetimes
3. Relativistic Spacetimes

### II. Classical and Relativistic Spacetimes.

Motivation: Turing solvable problems require a TM to halt after a finite number of steps with a given output.

- What if we allow TMs to perform an *infinite* number of steps?
  - Then: *Some Turing unsolvable problems may become solvable!*
  - But: *How could puny finite humans access the output of a TM that performs an infinite number of steps?*

Answer to come: Place the puny human in a spacetime that:

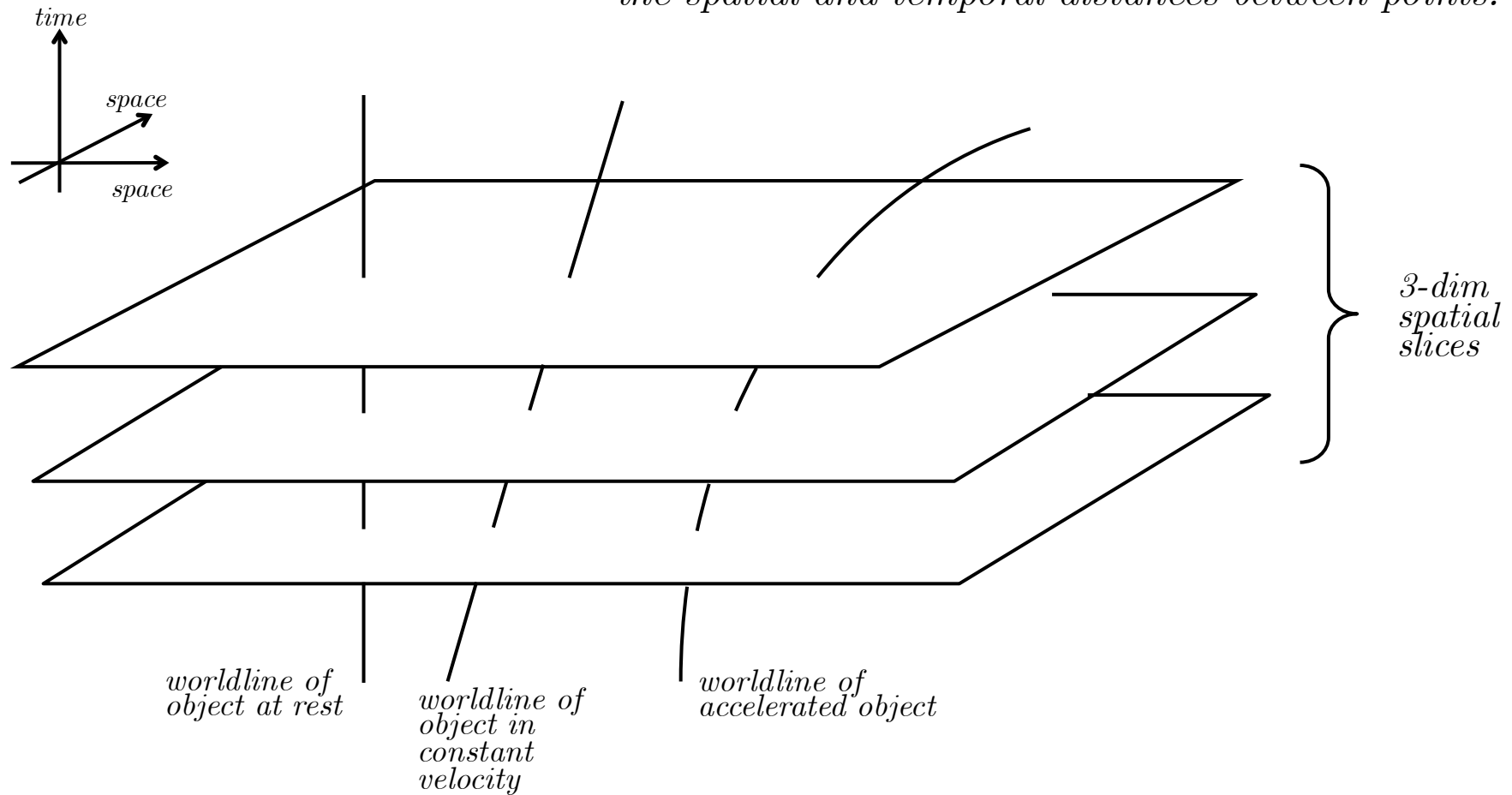
- (a) Allows the TM to live to infinity in its rest frame; and
- (b) Allows the human to access the TM's output in a finite amount of time.

- Mathematically: A matter of determining the appropriate curved geometry for the given spacetime.
- Physically: Are such appropriately curved spacetimes physically possible?

# 1. Types of Spacetimes

- A *spacetime* is a 4-dim collection of points with additional structure.

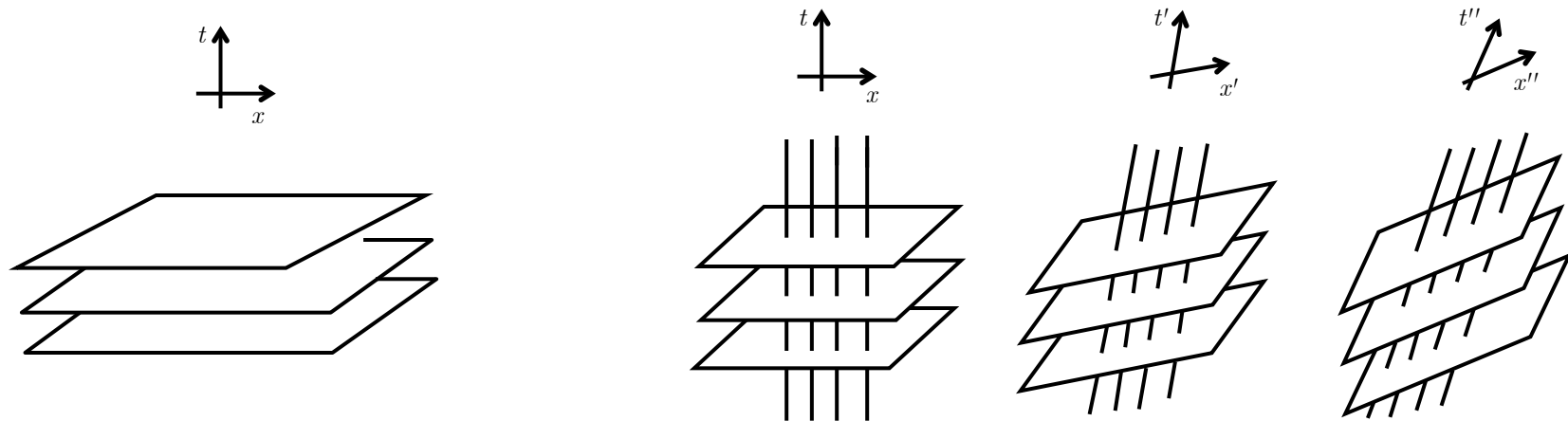
Typically, one or more metrics = a specification of the spatial and temporal distances between points.



Two ways spacetimes can differ:

(1) Different ways of specifying distances between points yield different types of spacetimes.

- *Classical spacetimes* have *separate* spatial and temporal metrics: only one way to split time from space (spatial and temporal distances are *absolute*).
- *Relativistic spacetimes* have a *single* spatiotemporal metric, and how it gets split into spatial and temporal parts depends on one's inertial reference frame (spatial and temporal distances are *relative*).



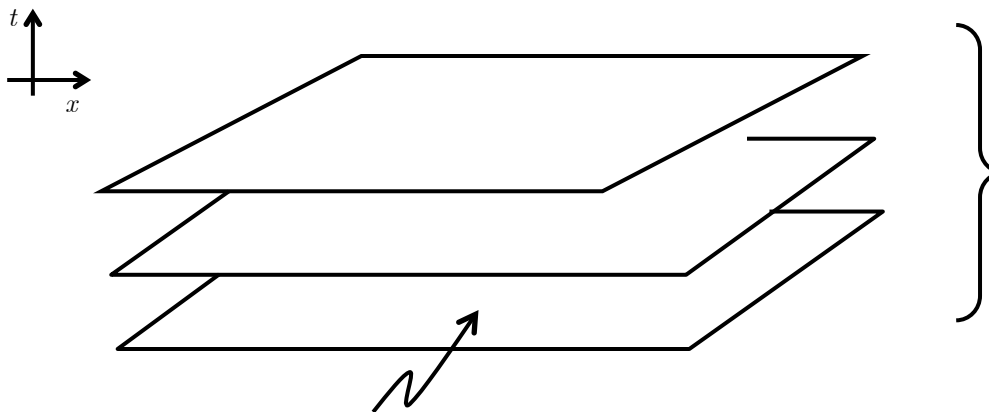
*Classical spacetime: only one way to split time from space.*

*Relativistic spacetime: many ways to split time from space.*

(2) Metrics can be *flat* or *curved*: how one specifies the distance between points encodes the curvature of the spacetime.

- *Classical spacetimes* can be flat or curved.
- *Relativistic spacetimes* can be flat (Minkowski spacetime) or curved (general relativistic spacetimes).

• Two ways curvature can manifest itself:



*How the spatial slices are "rigged" together can be flat or curved.*

*The spatial slices can be flat or curved.*

## 2. Classical Spacetimes

Newtonian spacetime is a 4-dim collection of points such that:

(N1) Between any two points  $p, q$  with coordinates  $(t, x, y, z)$  and  $(t + \Delta t, x + \Delta x, y + \Delta y, z + \Delta z)$  there is a definite *temporal interval*  $T(p, q) = \Delta t$ .

(N2) Between any two points  $p, q$  with coordinates  $(t, x, y, z)$  and  $(t + \Delta t, x + \Delta x, y + \Delta y, z + \Delta z)$  there is a definite *Euclidean distance*

$$R(p, q) = \sqrt{(\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2}$$

(N1) and (N2) entail:

(a) All worldlines have a definite *absolute velocity*.

For worldline  $\gamma$ , and any two points  $p, q$  on  $\gamma$ , the *absolute velocity* of  $\gamma$  with respect to  $p, q$  can be defined by  $R(p, q)/T(p, q)$ .

(b) There is a privileged collection of worldlines defined by  $R(p, q)/T(p, q) = 0$ .

(c) All worldlines have a definite *absolute acceleration*.

For worldline  $\gamma$ , and points  $p, q$  on  $\gamma$ , the *absolute acceleration* of  $\gamma$  with respect to  $p, q$  can be defined by  $d/dt\{R(p, q)/T(p, q)\}$ .

↖ Newton's  
absolute  
space!

***But absolute space and absolute velocity are unobservable!***

Neo-Newtonian spacetime is a 4-dim collection of points such that:

(NN1) Between any two points  $p, q$  with coordinates  $(t, x, y, z)$  and  $(t + \Delta t, x + \Delta x, y + \Delta y, z + \Delta z)$  there is a definite *temporal interval*  $T(p, q) = \Delta t$ .

(NN2) Between any two *simultaneous* points  $p_t, q_t$  with coordinates  $(t, x, y, z)$  and  $(t, x + \Delta x, y + \Delta y, z + \Delta z)$  there is a definite *Euclidean distance*,

$$R(p_t, q_t) = \sqrt{(\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2}$$

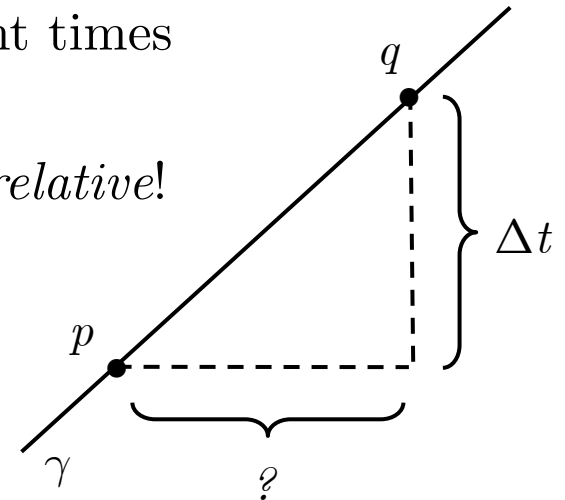
(NN3) Any worldline  $\gamma$  through a point  $p$  has a definite curvature  $S(\gamma, p)$ .

(NN2) entails:

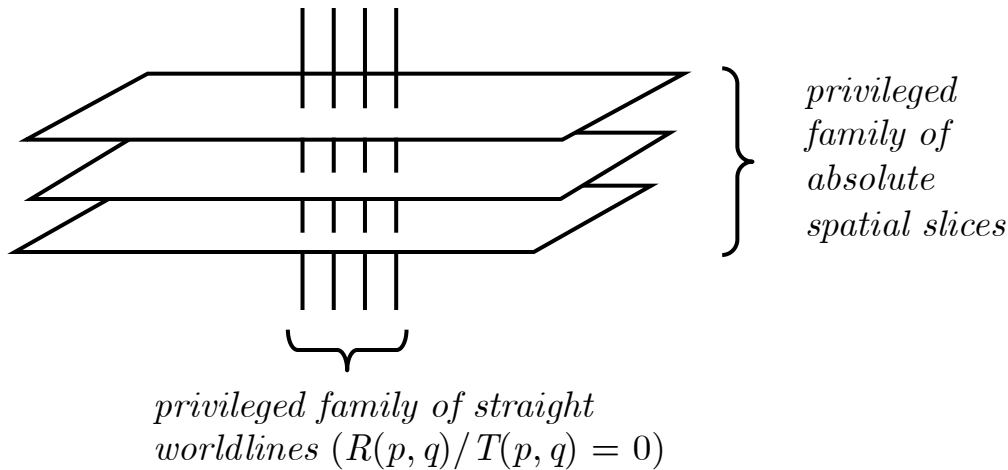
- No absolute spatial distance between points at different times on any worldline  $\gamma$ .
- So: No absolute velocity for any worldline: *velocity is relative!*
- So: No single privileged inertial reference frame.

(NN3) entails: acceleration remains absolute!

For worldline  $\gamma$  and point  $p$  on  $\gamma$ , the *absolute acceleration of  $\gamma$  with respect to  $p$*  is given by  $S(\gamma, p)$ .

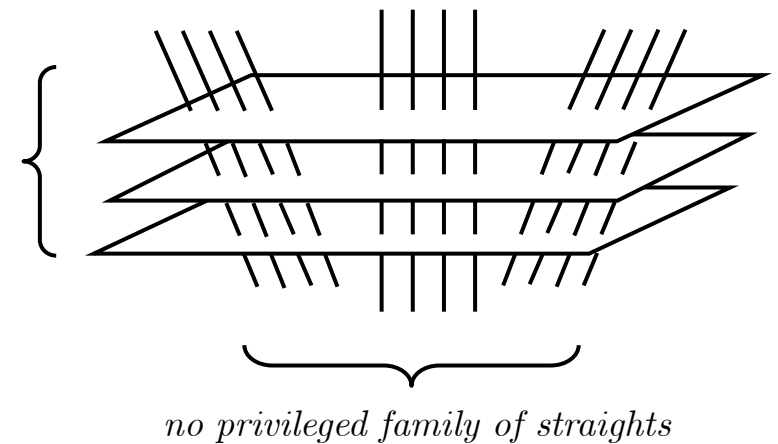


## Newtonian Spacetime



1. Single, privileged inertial frame.
2. Velocity is absolute.
3. Acceleration is absolute.
4. Simultaneity is absolute.

## Neo-Newtonian Spacetime



1. Many inertial frames; none privileged.
2. Velocity is relative.
3. Acceleration is absolute.
4. Simultaneity is absolute.

- Both Newtonian spacetime and Neo-Newtonian spacetime have absolute temporal metrics: *Everyone agrees on what time it is.*
- Relativistic spacetimes have no absolute temporal metric: *What time it is depends on your inertial reference frame.*

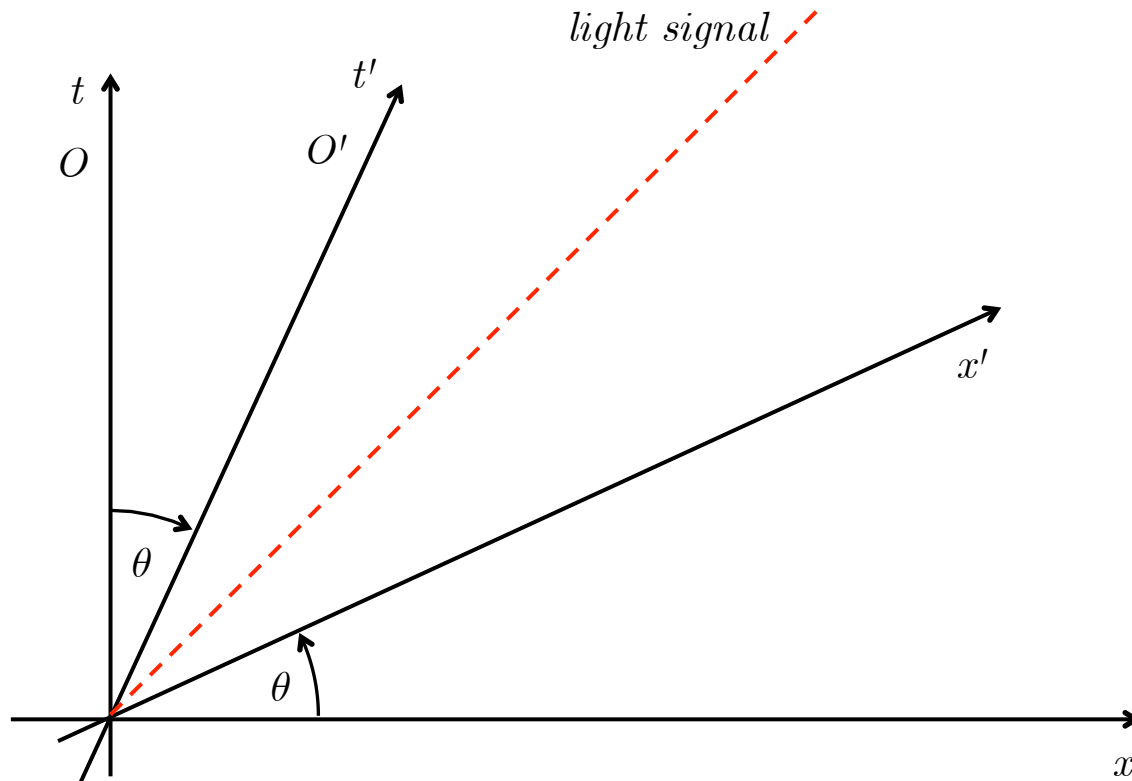
### 3. Relativistic Spacetimes

- Light Postulate of Special Relativity entails:

*The speed of light  $c$  is the same in all inertial reference frames.*



*Albert Einstein*  
(1879-1955)



$$c = \frac{\Delta x}{\Delta t} = \frac{\Delta x'}{\Delta t'}$$

*value of  $c$  for  $O$       value of  $c$  for  $O'$*

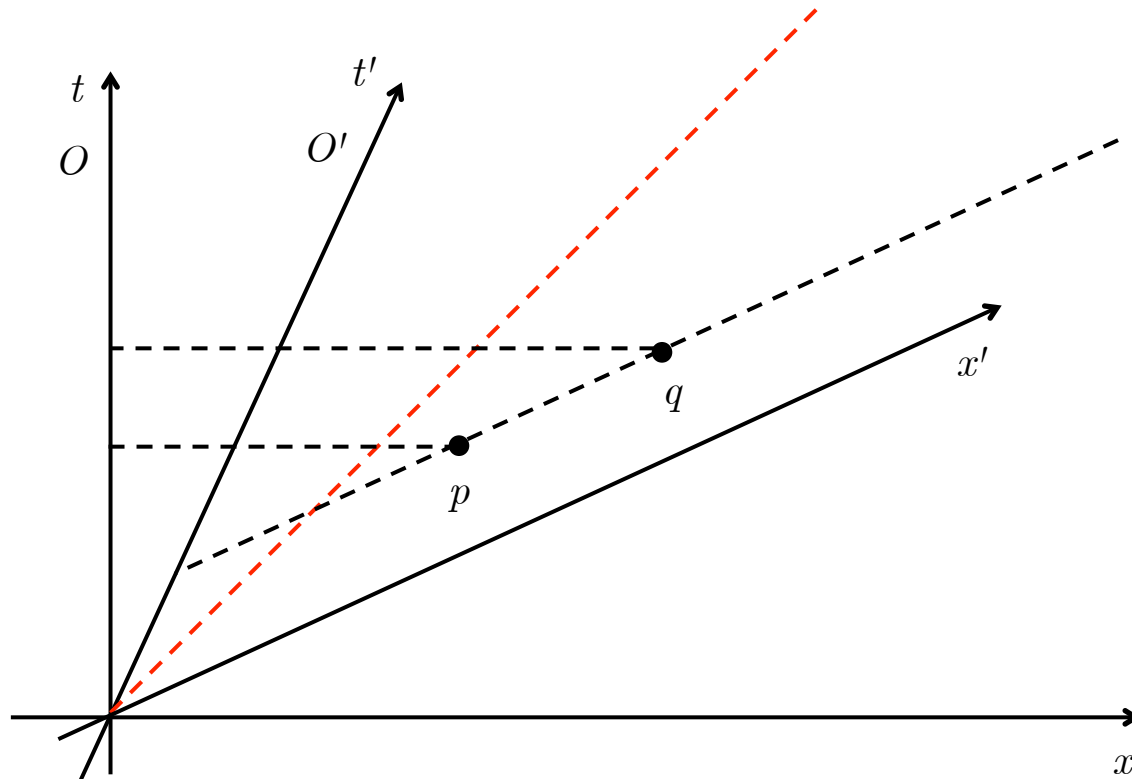
- Object  $O'$  is moving at constant velocity with respect to object  $O$ .
- $O$  and  $O'$  must measure same speed  $c$  for light signal.
- So: The  $x'$  axis must be inclined by the same amount  $\theta$  from the  $x$  axis as the  $t'$  axis is inclined from the  $t$  axis.
- Thus:  $O$  and  $O'$  must disagree on spatial and temporal measurements!



### 3. Relativistic Spacetimes

- Light Postulate of Special Relativity entails:

*The speed of light  $c$  is the same in all inertial reference frames.*



- $O$  and  $O'$  make different judgements of simultaneity (*relativity of simultaneity*).
- $p$  and  $q$  are simultaneous according to  $O'$ .
- $p$  happens before  $q$  according to  $O$ .

## Spacetime of Special Relativity = Minkowski spacetime

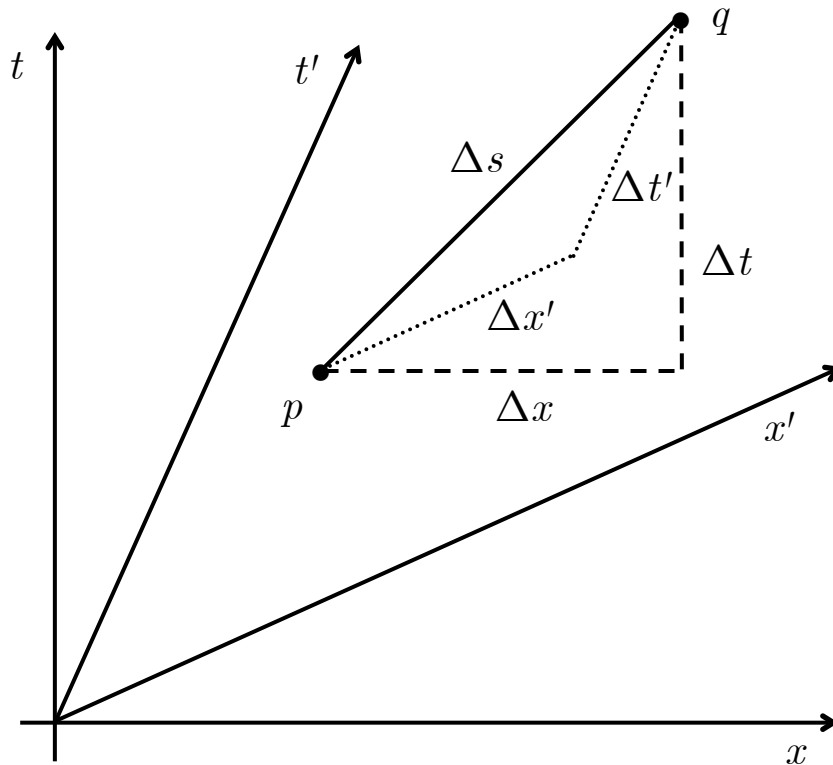


Hermann  
Minkowski  
(1864-1909)

Minkowski spacetime = 4-dim collection of points such that between any two points  $p$ ,  $q$  with coordinates  $(t, x, y, z)$  and  $(t + \Delta t, x + \Delta x, y + \Delta y, z + \Delta z)$  there is a definite spacetime interval given by

$$\Delta s = \sqrt{-(c\Delta t)^2 + (\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2}.$$

- Similar to Euclidean *spatial* interval  $\sqrt{(\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2}$ .
- But: Includes the time coordinate difference, too! And it's *negative*!
  - Idea: All inertial frames will agree on the spatiotemporal distance  $\Delta s$  between any points  $p$  and  $q$ .
  - But they will disagree on how  $\Delta s$  gets split into a temporal part and a spatial part: they will disagree on measurements of time and measurements of space.



$$\begin{aligned}\Delta s &= \sqrt{-(c\Delta t)^2 + (\Delta x)^2} \\ &= \sqrt{-(c\Delta t')^2 + (\Delta x')^2}\end{aligned}$$

- All inertial frames agree on the *spacetime* distance between any two points  $p$  and  $q$ .
- They will disagree on the *temporal* distance between  $p$  and  $q$  (time dilation) and on the *spatial* distance (length contraction).
  - They will disagree on how they split the spacetime distance into temporal and spatial parts.

The Minkowski spacetime interval is encoded in the *Minkowski metric*  $\eta_{\mu\nu}$

$$\begin{aligned}
 (\Delta s)^2 &= -(c\Delta t)^2 + (\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2 \\
 &= \sum_{\mu,\nu=0}^3 \eta_{\mu\nu} \Delta x^\mu \Delta x^\nu
 \end{aligned}$$

$$\begin{aligned}
 \Delta x^0 &= c\Delta t, \quad \Delta x^1 = \Delta x, \\
 \Delta x^2 &= \Delta y, \quad \Delta x^3 = \Delta z, \\
 \eta_{\mu\nu} &= \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}
 \end{aligned}$$

- Infinitesimally:  $ds^2 = \eta_{\mu\nu} dx^\mu dx^\nu$

Three forms of  $(\Delta s)^2$

(a) *Timelike*.  $(\Delta s)^2 < 0$ , or:  $\frac{\sqrt{(\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2}}{\Delta t} < c$

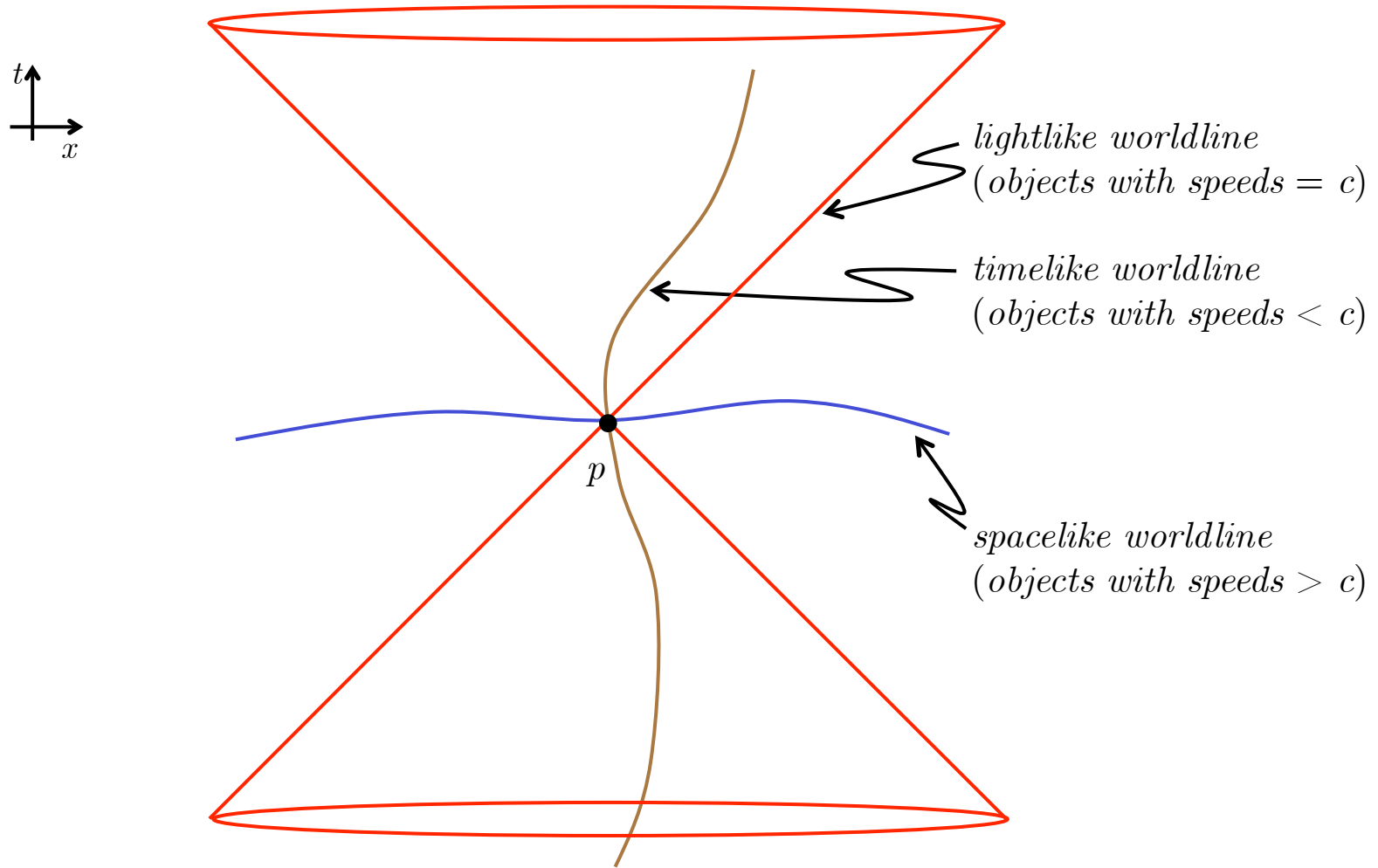
(b) *Lightlike*.  $(\Delta s)^2 = 0$ , or:  $\frac{\sqrt{(\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2}}{\Delta t} = c$

(c) *Spacelike*:  $(\Delta s)^2 > 0$ , or:  $\frac{\sqrt{(\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2}}{\Delta t} > c$

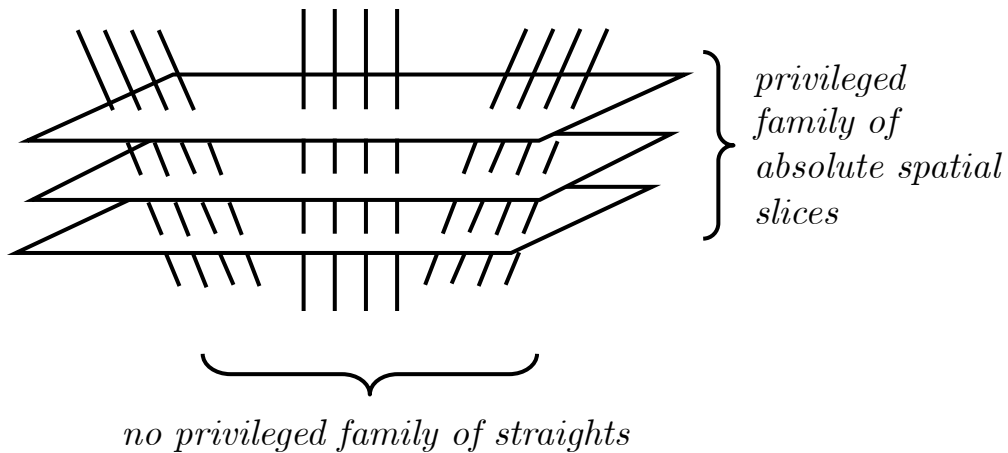
↪ *Three different types of worldline in Minkowski spacetime!*

- Absolute distinction: All inertial frames agree on  $\Delta s$ , so all inertial frames agree on which worldlines are timelike, lightlike, and spacelike!

Hence: The Minkowski metric defines a *lightcone* at any point  $p$ :

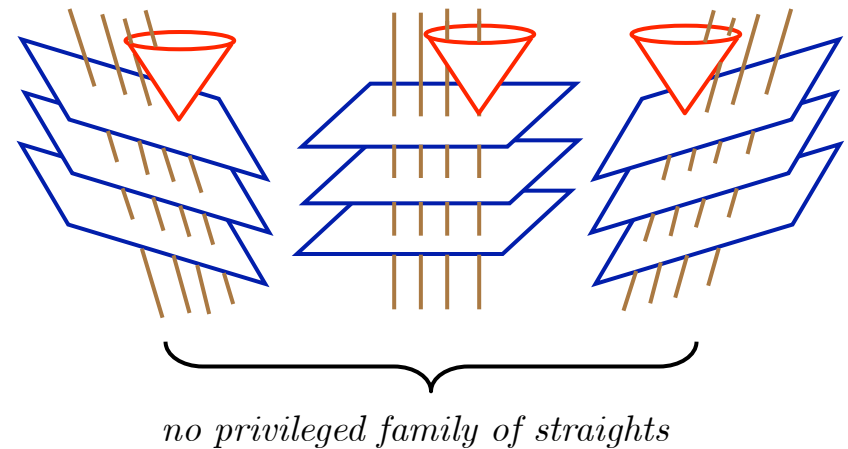


## Neo-Newtonian Spacetime



1. Many inertial frames; none privileged.
2. Velocity is relative.
3. Acceleration is absolute.
4. Simultaneity is absolute.

## Minkowski Spacetime



1. Many inertial frames; none privileged.
2. Velocity is relative.
3. Acceleration is absolute.
4. Simultaneity is relative.
5. Invariant light-cone structure at each point.

Problem: Special relativity does not account for the gravitational force.

- To include gravity...

*Geometricize it! Make it a feature of spacetime geometry.*



Two requirements:

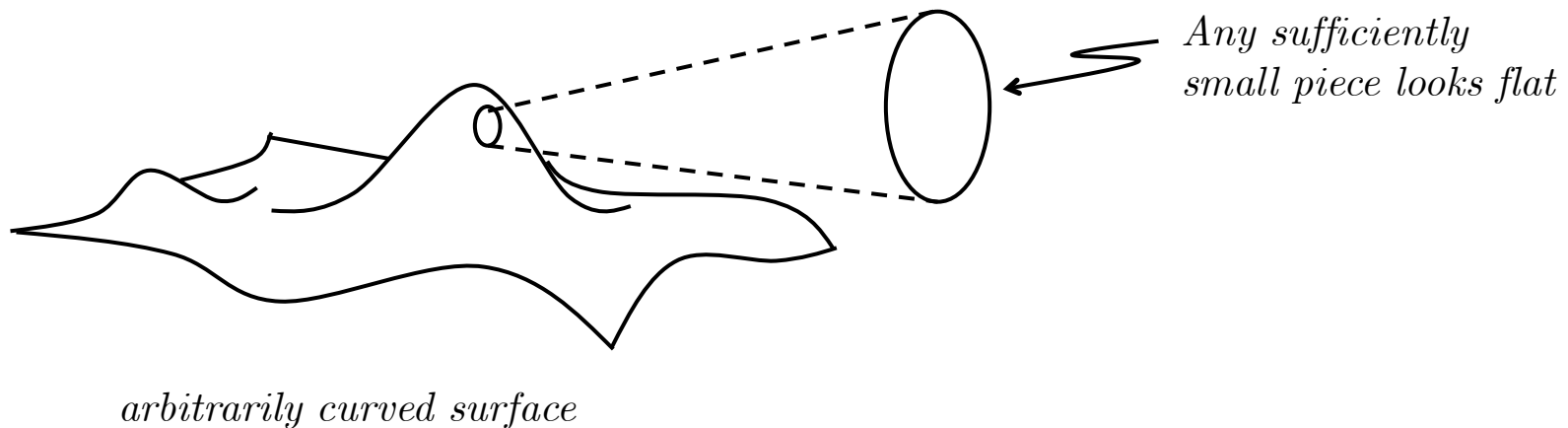
(1) New theory ("general relativity") must reduce to special relativity in sufficiently flat regions of spacetime:

- Replace  $ds^2 = \eta_{\mu\nu} dx^\mu dx^\nu$  with  $ds^2 = g_{\mu\nu} dx^\mu dx^\nu$ .

*flat Minkowski metric*

*non-flat metric*

- Require  $g_{\mu\nu}$  to reduce to  $\eta_{\mu\nu}$  in small regions of spacetime.



Problem: Special relativity does not account for the gravitational force.

- To include gravity...

*Geometricize it! Make it a feature of spacetime geometry.*



Two requirements:

(2) Curvature of spacetime must be related to matter density:

- The Einstein equations (1916):

$$G_{\mu\nu} = \kappa T_{\mu\nu}$$


*Einstein tensor = encodes curvature of spacetime as a function of  $g_{\mu\nu}$*

*Stress-energy tensor = encodes matter density*

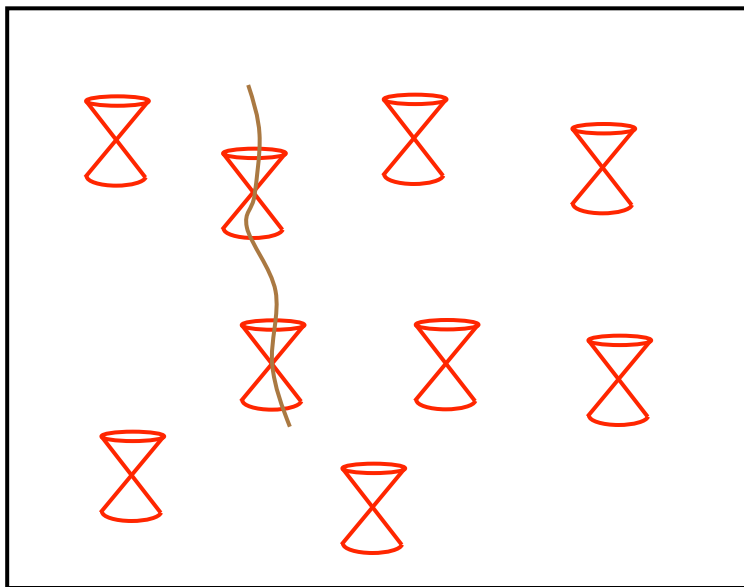
- Consequence: The Minkowski metric is the solution for zero curvature  $G_{\mu\nu} = 0$  (*i.e.*, spatiotemporal flatness).



A general relativistic spacetime = 4-dim collection of points such that between any two (infinitesimally close) points, there is a definite spacetime interval given by  $ds^2 = g_{\mu\nu} dx^\mu dx^\nu$ , where  $g_{\mu\nu}$  is a Lorentzian metric that satisfies the Einstein equations.

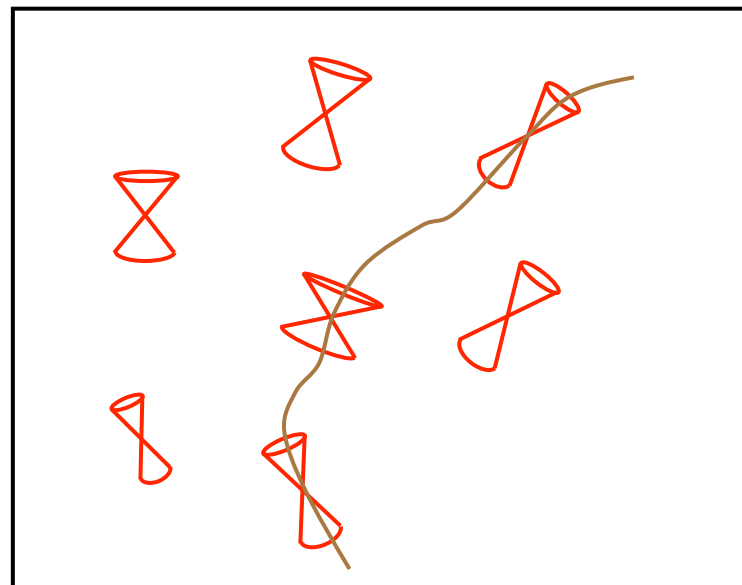
  
"reduces to the Minkowski metric at any point"

## Minkowski Spacetime



Invariant light-cone structure at each point: light-cones all have same size and orientation.

## Arbitrary General Relativistic Spacetime



Light-cone structure at each point is not invariant: light-cones may not have same size and orientation due to curvature.

- Idea: The light-cone structure constrains the motion of physical objects (traveling on timelike worldlines).
- And: In an arbitrary general relativistic spacetime, the matter density determines the metric, which determines the light-cone structure.