

16. Black Hole Thermodynamics. Part 1.

1. Gravitational Collapse.

Gravitational Collapse in Newtonian Physics

- Michell (1784) "On the Means of discovering the Distance, Magnitude, *etc.* of the Fixed Stars...", *Phil. Trans. Roy. Soc. London*, lxxiv, 35-57.
- Escape velocity of sun = 497 times smaller than velocity of light.
- So: Any object with same density as sun and radius 497 times larger will trap all light that falls on it:

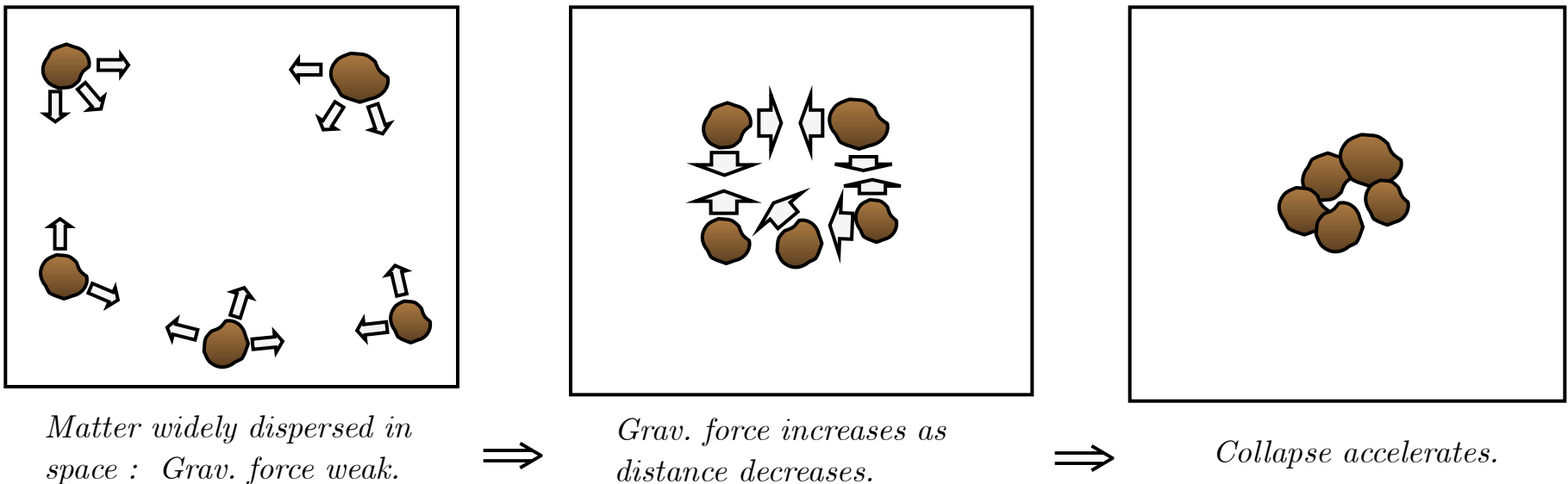


John Michell
(1724-1793)

"... if the semi-diameter of a sphere of the same density with the sun were to exceed that of the sun in the proportion of 500 to 1, a body falling from an infinite height towards it would have acquired at its surface a greater velocity than that of light, and consequently, supposing light to be attracted by the same force in proportion to its *vis inertiae*, with other bodies, all light emitted from such a body would be made to return towards it, by its own proper gravity."

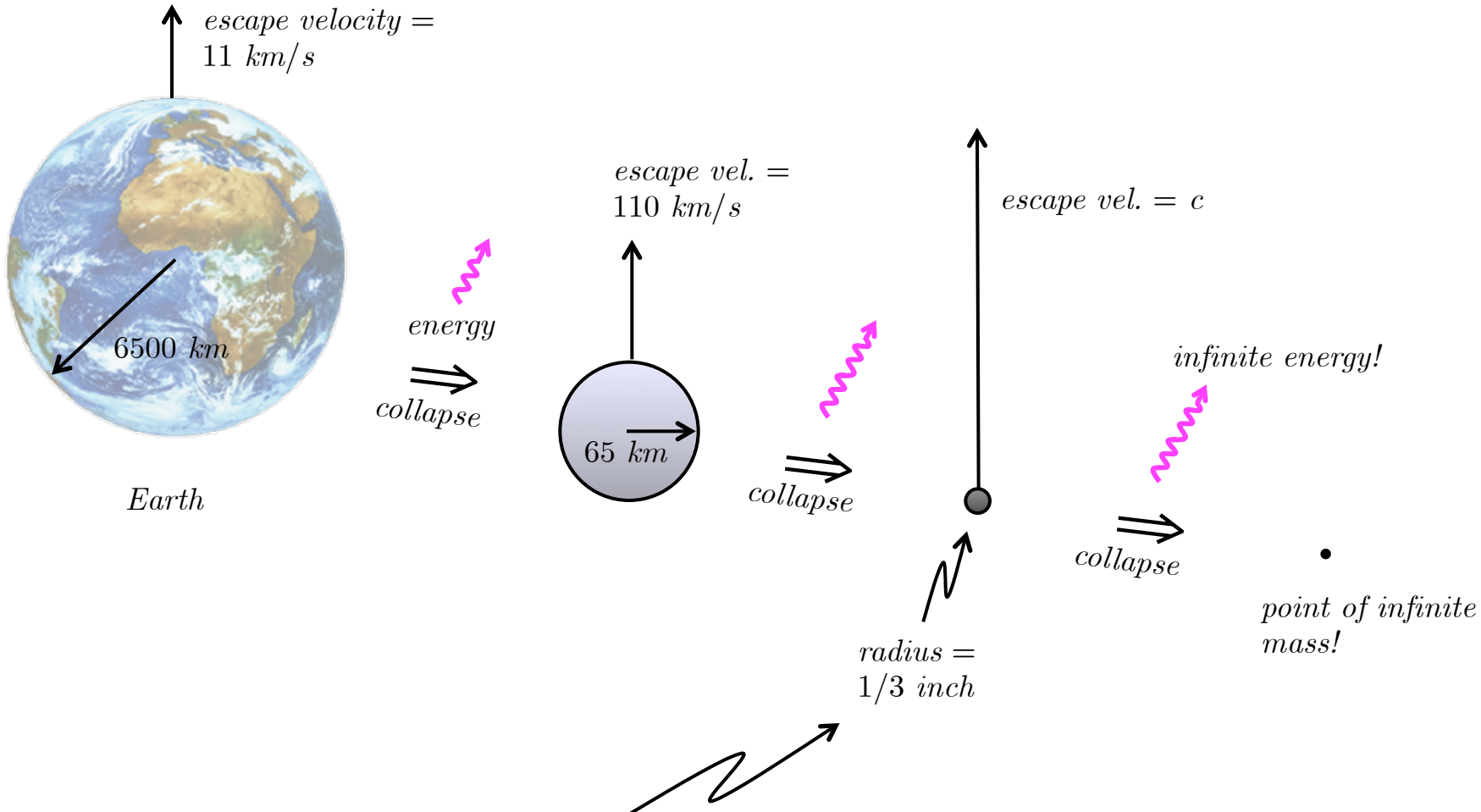
- How might such objects originate?

- Consider Newtonian gravitational collapse: $F_G \propto 1/r^2$.
 - Decreasing the distance by half increases the force by 4.



- Is this an entropy-decreasing process?
 - For *interaction-free* gas, increased "clumping" is associated with decrease in entropy.
 - But: For *self-gravitating* gas, decrease in entropy due to clumping is typically overcome by increase in entropy due to *release of energy*.

- In Newtonian gravitational collapse, if nothing intervenes, matter collapses to a point with *infinite* energy release.



At this stage, light cannot escape surface: Earth is now invisible to outside observers (Newtonian black hole).

Newtonian Picture: Complete Gravitational Collapse

$$\left[\begin{array}{c} \textit{Unrestrained} \\ \textit{gravitational} \\ \textit{collapse} \end{array} \right] \Rightarrow \left[\begin{array}{c} \textit{infinite} \\ \textit{energy} \\ \textit{release} \end{array} \right] + \left[\begin{array}{c} \textit{singular point} \\ \textit{of infinite} \\ \textit{matter density} \end{array} \right] + \left[\begin{array}{c} \textit{background} \\ \textit{spacetime} \\ \textit{unaffected} \end{array} \right]$$

In General Relativity (Einstein 1916):

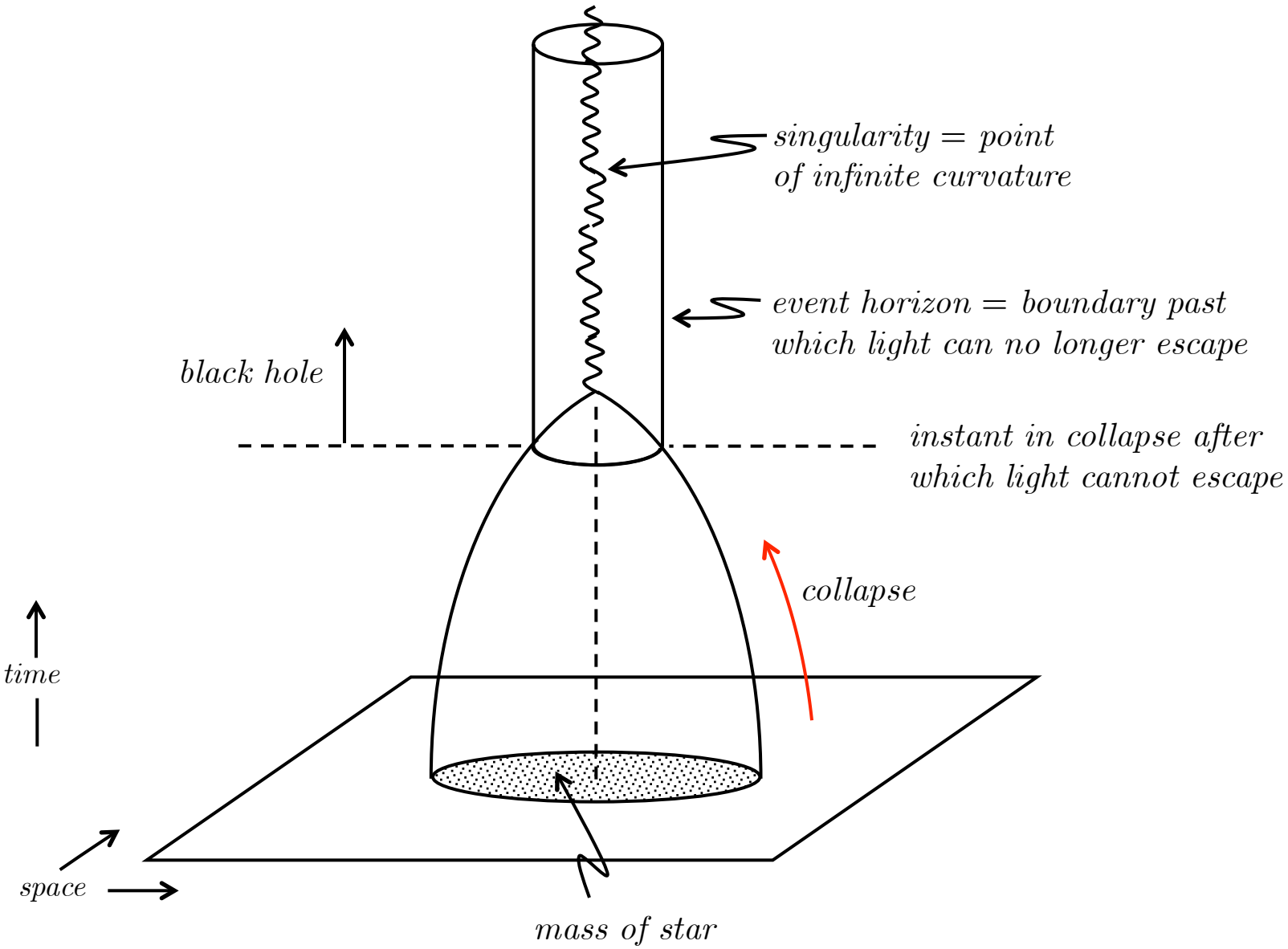
- Equivalence of mass and energy.
- Curvature of spacetime is determined by mass-energy density.

Relativistic Picture: Complete Gravitational Collapse

$$\left[\begin{array}{c} \textit{Unrestrained} \\ \textit{gravitational} \\ \textit{collapse} \end{array} \right] \Rightarrow \left[\begin{array}{c} \textit{finite energy} \\ \textit{release} \end{array} \right] + \left[\begin{array}{c} \textit{singular point of infinite} \\ \textit{spacetime curvature} \end{array} \right]$$

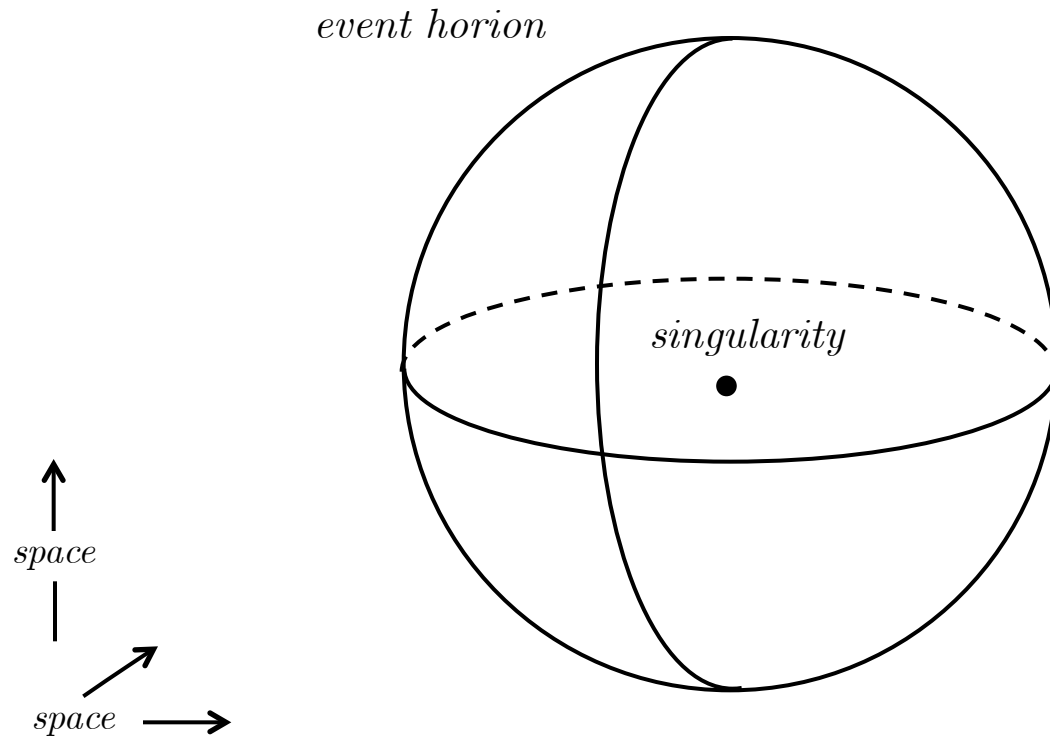
2. Relativistic Black Holes.

Ex 1: Schwarzschild black hole: no charge or rotation (Schwarzschild 1916).



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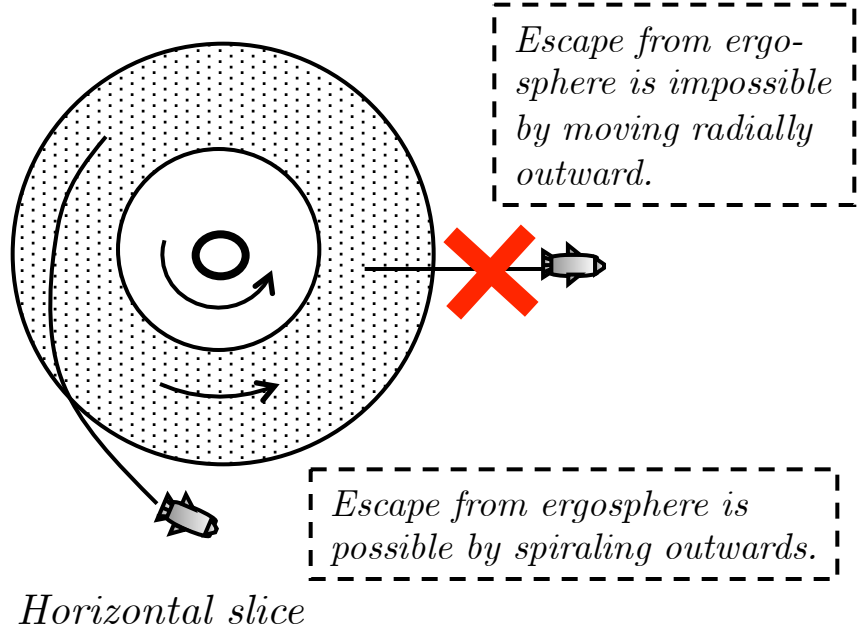
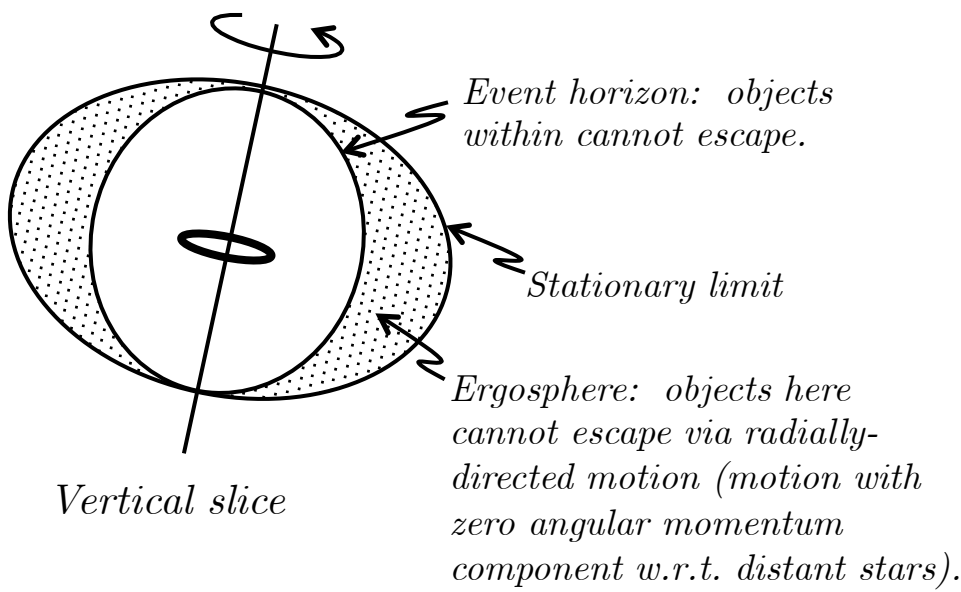
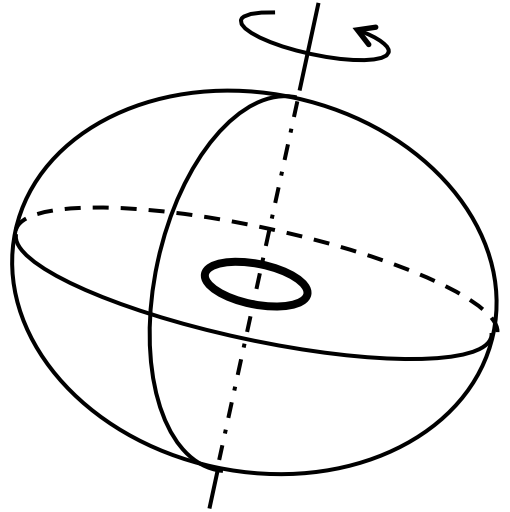
Ex 2: Rotating black hole. (Kerr 1963).

- Consequences of rotation:

- Singularity is a ring, not a point.
- Inertial frames near singularity are dragged.

- Stationary limit = distance from singularity at which objects can no longer have zero angular momentum with respect to distant stars.

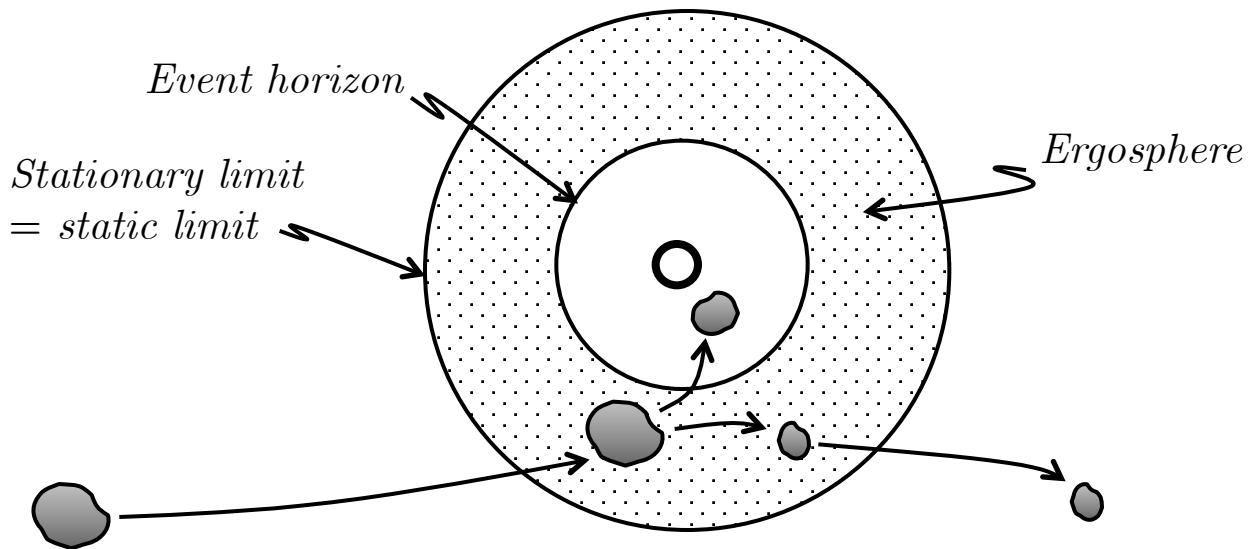
- Ergosphere = Region between event horizon and stationary limit.



- Static limit = distance from singularity at which objects can no longer escape *via* radially directed motion.
- Inside static limit, the "Killing" vector field ξ^a that encodes time-translation symmetries is *spacelike*: $|\xi^a| > 0$.
- So: Inside static limit, objects with *negative energy*, $E = -\xi^a p_a$, can exist.

- Allows the following procedure for extracting energy (Penrose 1969):

Negative energy object = object that requires more energy than its rest mass in order to move it to "infinity".



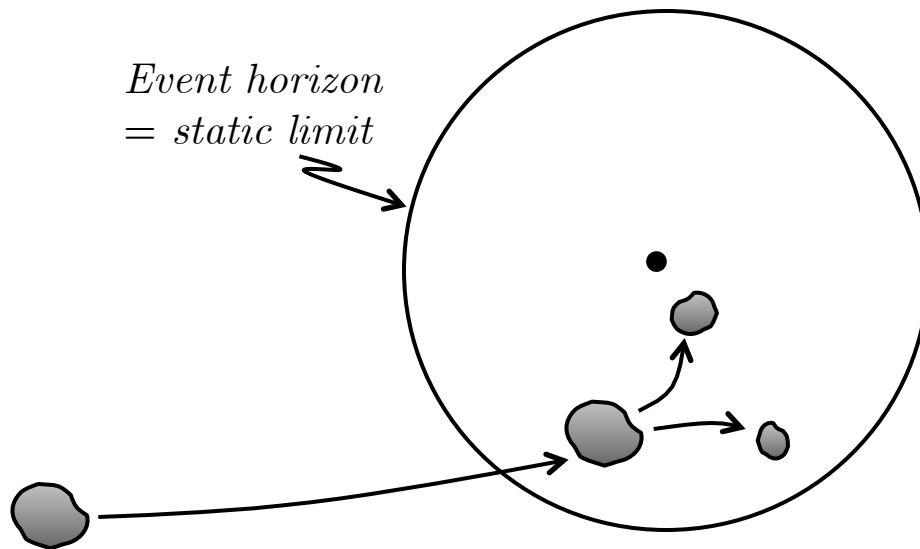
Lump of matter with energy E shot into ergosphere.

Fragments into two chunks with energies E_1, E_2 such that
 (a) $E_1 + E_2 = E$.
 (b) $E_1 < 0, E_2 > 0$.

Chunk with neg energy falls into singularity, chunk with positive energy emerges with greater energy than initial lump!

Why won't this work for a Schwarzschild black hole?

- Schwarzschild black hole: static limit = event horizon.
- So: Negative energy states are possible inside event horizon.
- But: To guarantee your rock is split into negative energy and positive energy pieces, you need to be inside the region of negative energy states when you split it!



*Event horizon
= static limit*

No ergosphere!

*Lump of matter
with energy E shot
into event horizon.*

*Fragments into two chunks
with energies E_1, E_2 such that*
(a) $E_1 + E_2 = E$.
(b) $E_1 < 0, E_2 > 0$.

Both chunks can't escape!

General Properties of Relativistic Black Holes

- No Hair Conjecture: A black hole is completely characterized by its mass M , charge Q , and angular momentum J .

Four types of black hole:

	<i>nonrotating ($J = 0$)</i>	<i>rotating ($J \neq 0$)</i>
<i>uncharged ($Q = 0$)</i>	Schwarzschild	Kerr
<i>charged ($Q \neq 0$)</i>	Reissner-Nordström	Kerr-Newman

- *Radius of event horizon*: $R_h = M + [M^2 - Q^2 - (J/M)^2]^{1/2}$
- *Area of event horizon*: $A = 4\pi[R_h^2 + (J/M)^2]$
 - So: A small change in mass δM will correspond to small changes in area δA , charge δQ , and angular momentum δJ .
- Hawking (1971) Area Theorem: $\delta A \geq 0$ in any process.
 - Ex: Suppose two black holes with areas A_1, A_2 collide to form black hole with area A_3 . Then $A_3 \geq A_1 + A_2$.



Stephen Hawking

3. The Laws of Black Hole Mechanics.

(Bardeen, Carter, Hawking 1973)

Black Hole Mechanics

Thermodynamics

0th Law

Surface gravity κ is constant over the event horizon of a stationary black hole.

κ is acceleration needed to keep an object at event horizon.

Temperature T is constant throughout a body in thermal equilibrium.

1st Law

$$\delta M = (1/8\pi)\kappa\delta A + \Phi\delta Q + \Omega\delta J$$

Φ is electrostatic potential, Ω is rotational velocity.

$$dE = TdS + pdV + \Omega dJ$$

2nd Law

$\delta A \geq 0$ in any process.

$\delta S \geq 0$ in any process.

3rd Law

$\kappa = 0$ is not achievable by any process.

$T = 0$ is not achievable by any process.

- Formally identical if $A/4 = S$ and $(1/2\pi)\kappa = T$.
- Is this merely a *formal equivalence*, or does it have a *physical basis*?